

Transformation of standing poloidal Alfven wave to toroidal Alfven wave due to the field line curvature

Dmitri Klimushkin,

Anatoli Leonovich,

Pavel Mager,

Institute of Solar-Terrestrial Physics, Irkutsk, Russia

Oleg Cheremnykh,

Space Research Institute, Kyiv, Ukraine

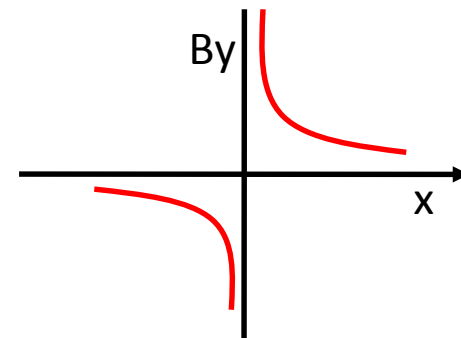
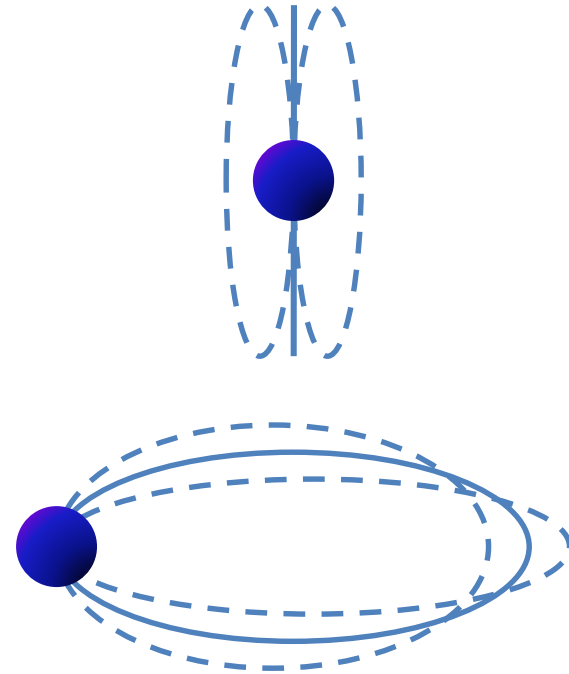
Prague, June 23, 2015

In memory of Vitali Mazur (1946-2015)

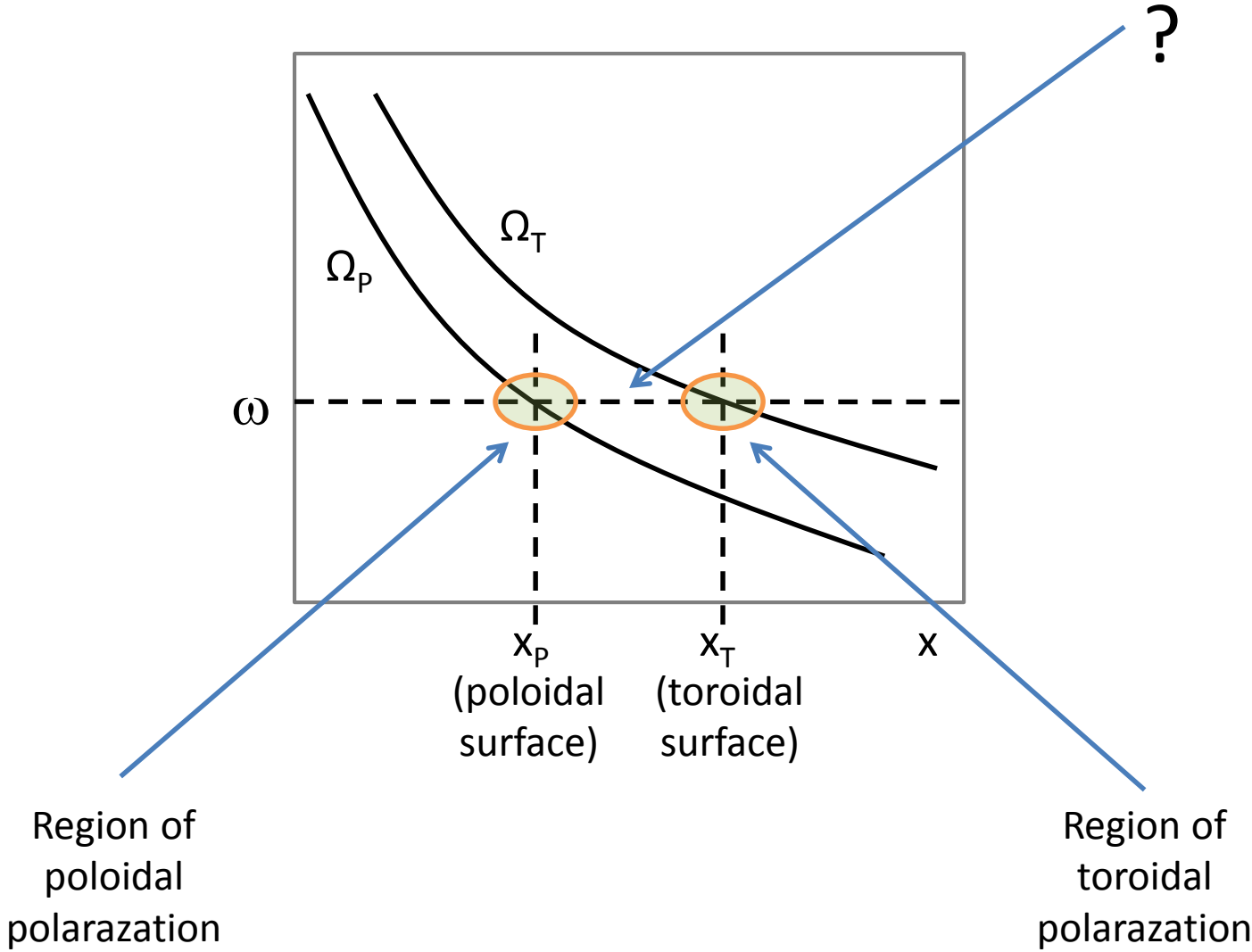


Three milestones

- 1954, J. Dungey, $m=0$:
 - Toroidal Alfvén modes ($B_y \gg B_x$)
- 1967, H.R. Radoski, $m \gg 1$:
 - Poloidal Alfvén modes ($B_x \gg B_y$);
 - different frequencies: $\Omega_T \neq \Omega_p$
- 1974: J. Southwood, L. Chen & A. Hasegawa, H.R. Radoski:
 - Logarithmic (pole) singularity at any m -number;
 - \rightarrow at any m -number the Alfvén wave can be toroidal ($B_y \gg B_x$)

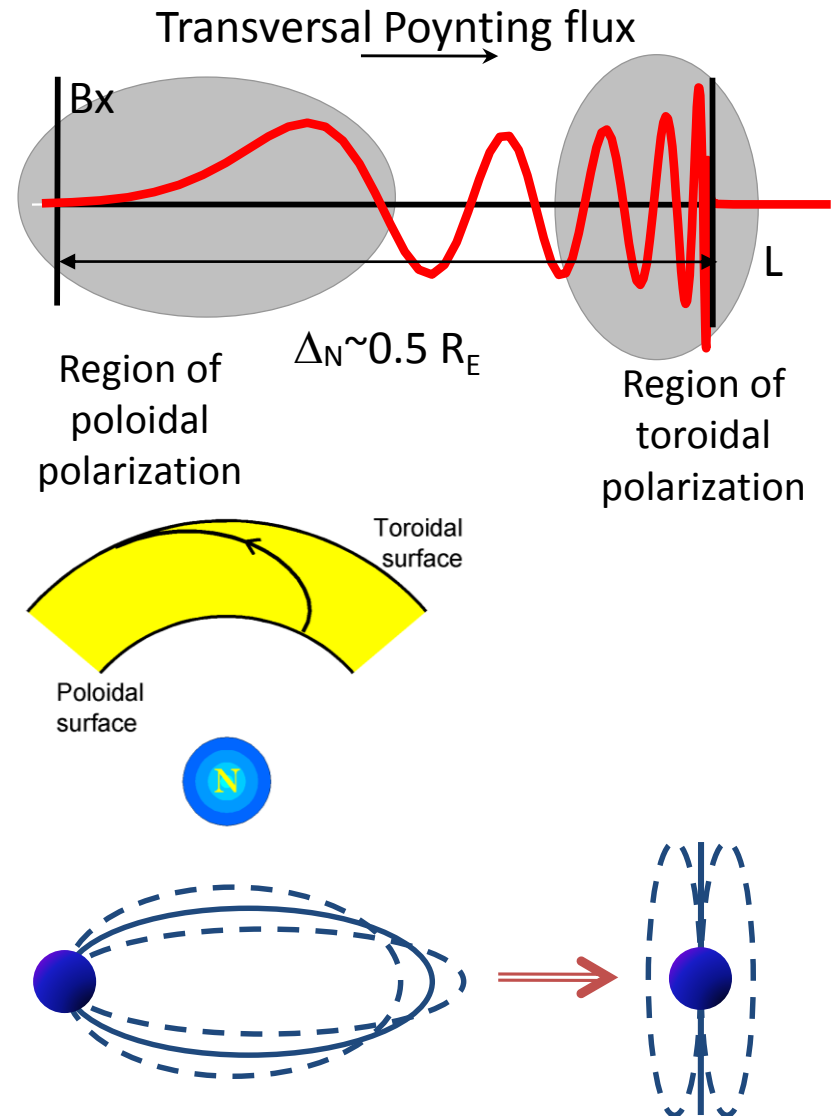


Graphical representation of $\Omega_T \neq \Omega_P$



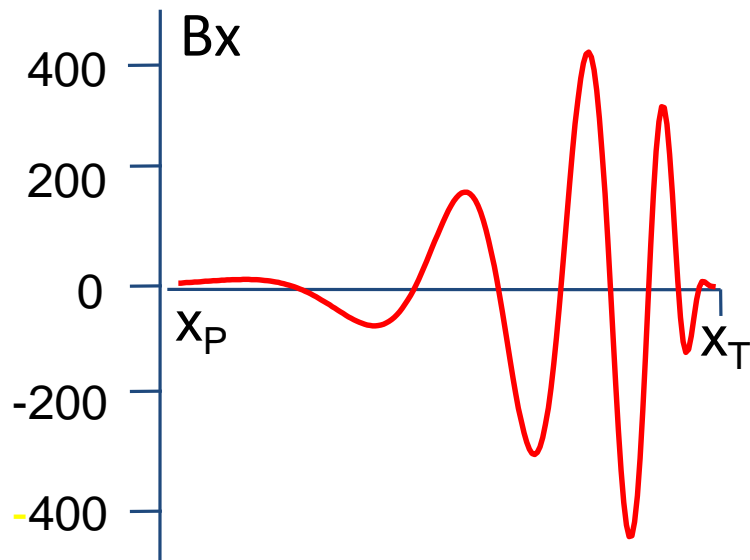
Global solution

- Leonovich & Mazur, 1993:
 - Wave travels from the poloidal to the toroidal surface;
 - The wave is standing along the field lines;
 - Radial group velocity and Poynting flux appear;
 - Beyond the region between poloidal and the toroidal surface: the mode is evanescent;
 - In the course of the propagation polarization changes from poloidal to toroidal;
 - Cause: field line curvature.

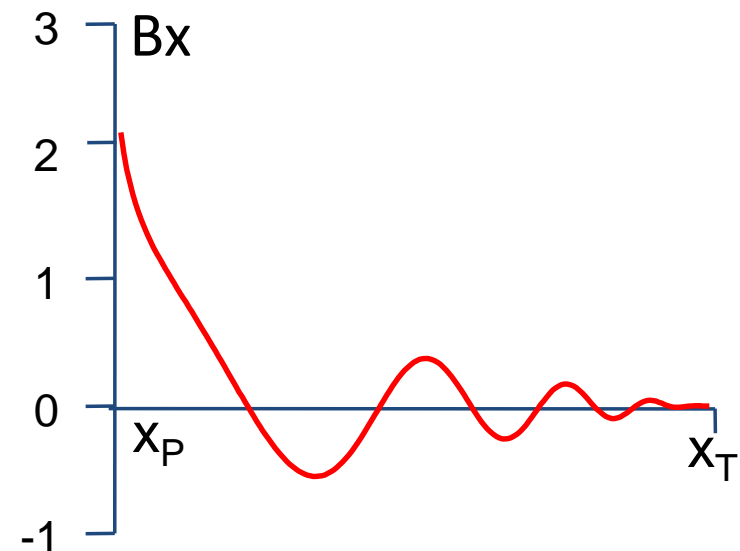


Instability due to the interaction with the energetic particles

- Klimushkin, 1998:
 - The wave is generated by some external current...
 - ... And is amplified by the particles in the course of the propagation across the magnetic shells



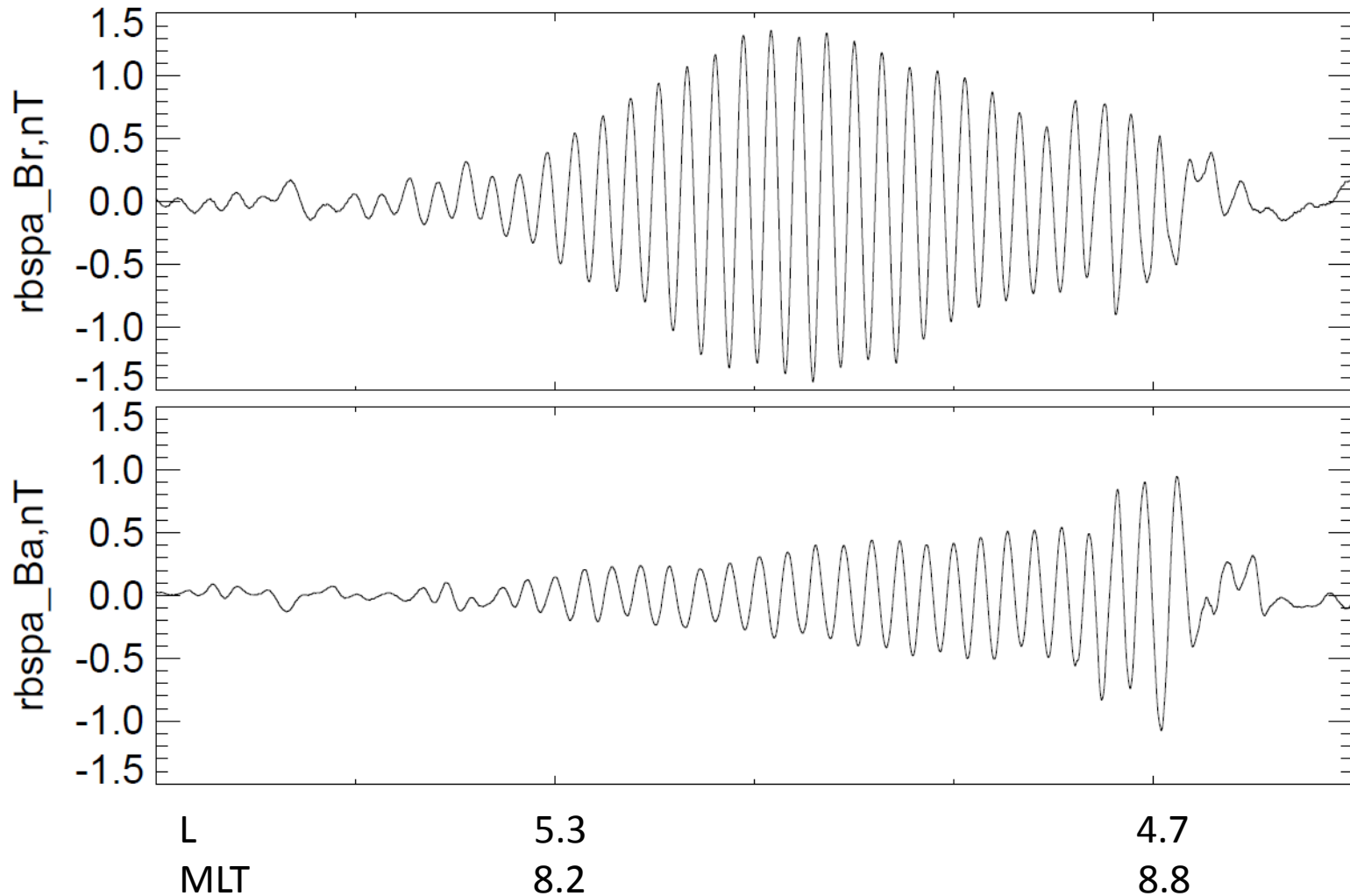
- Large instability, small ionospheric attenuation



- Large attenuation, small instability

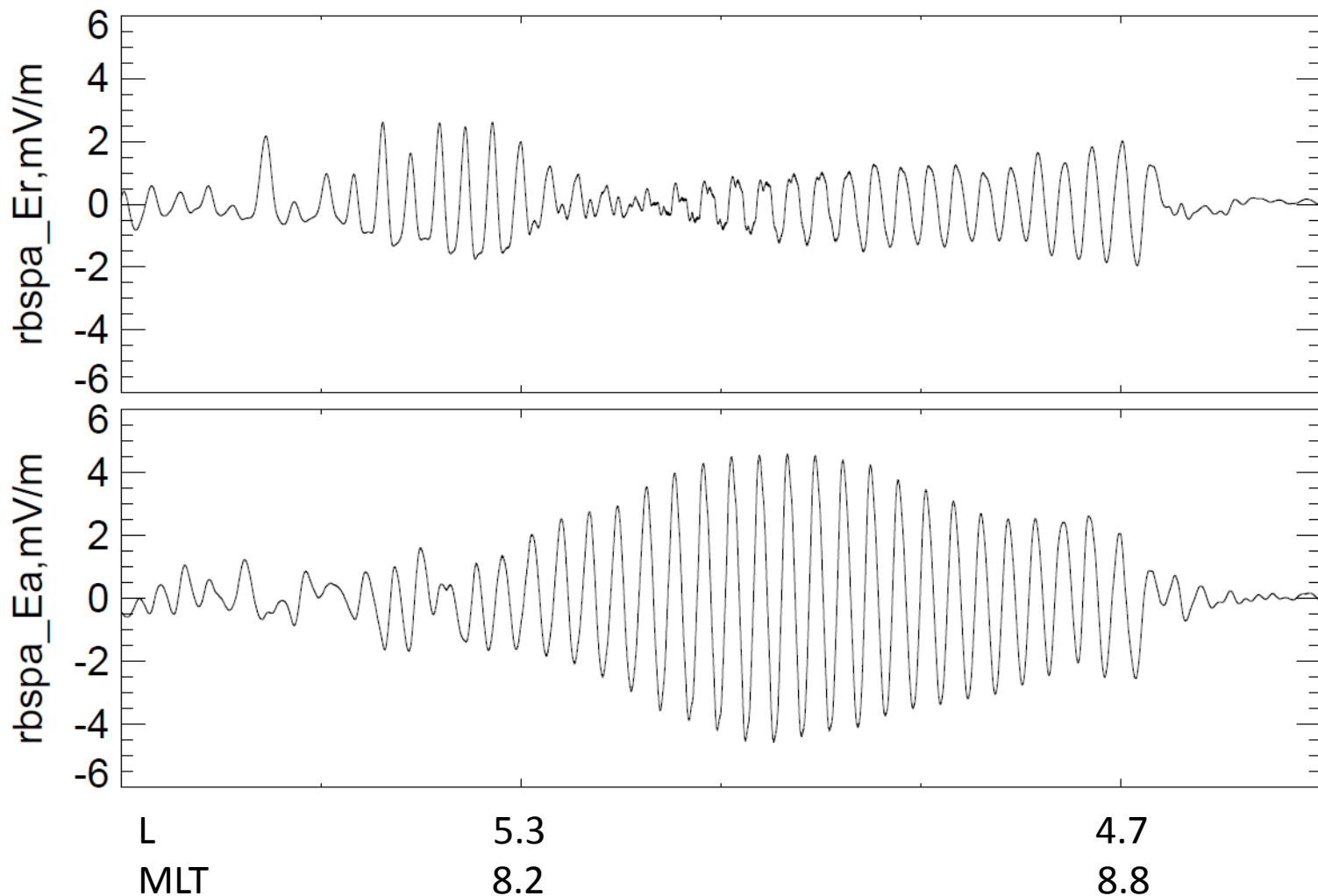
Lei Dai et al. 2013: Poloidal Alfvén wave generated by energetic particles - I

RBSP-A, Oct. 23, 2012. Magnetic field:

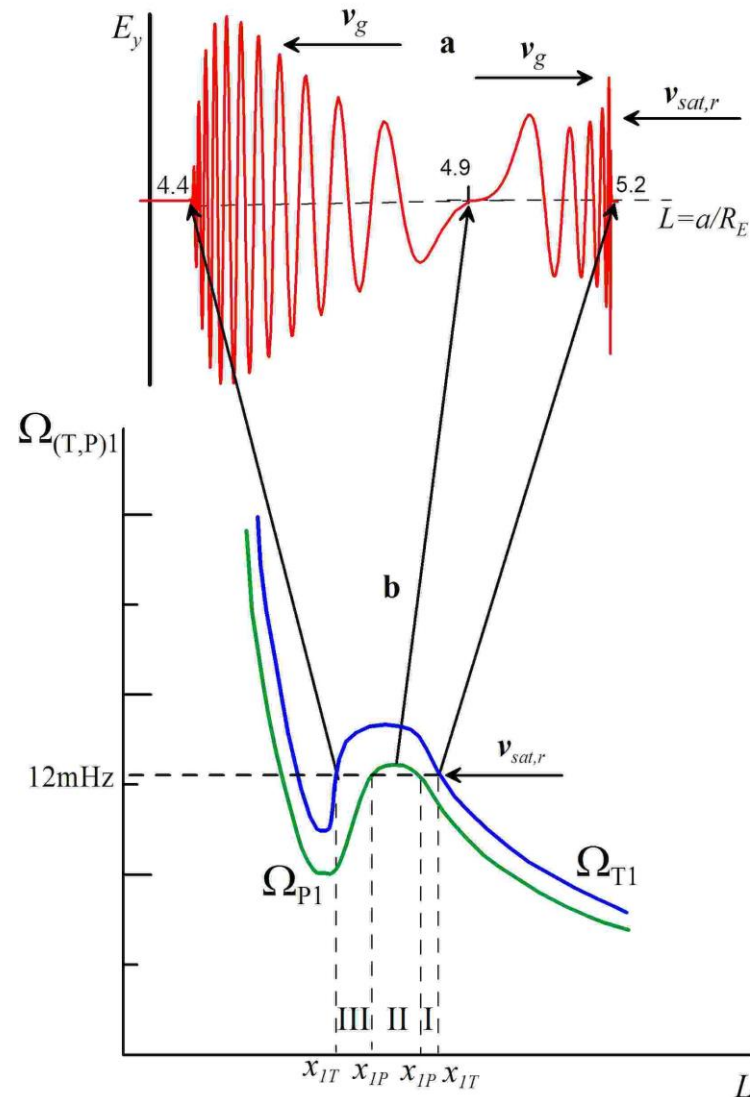
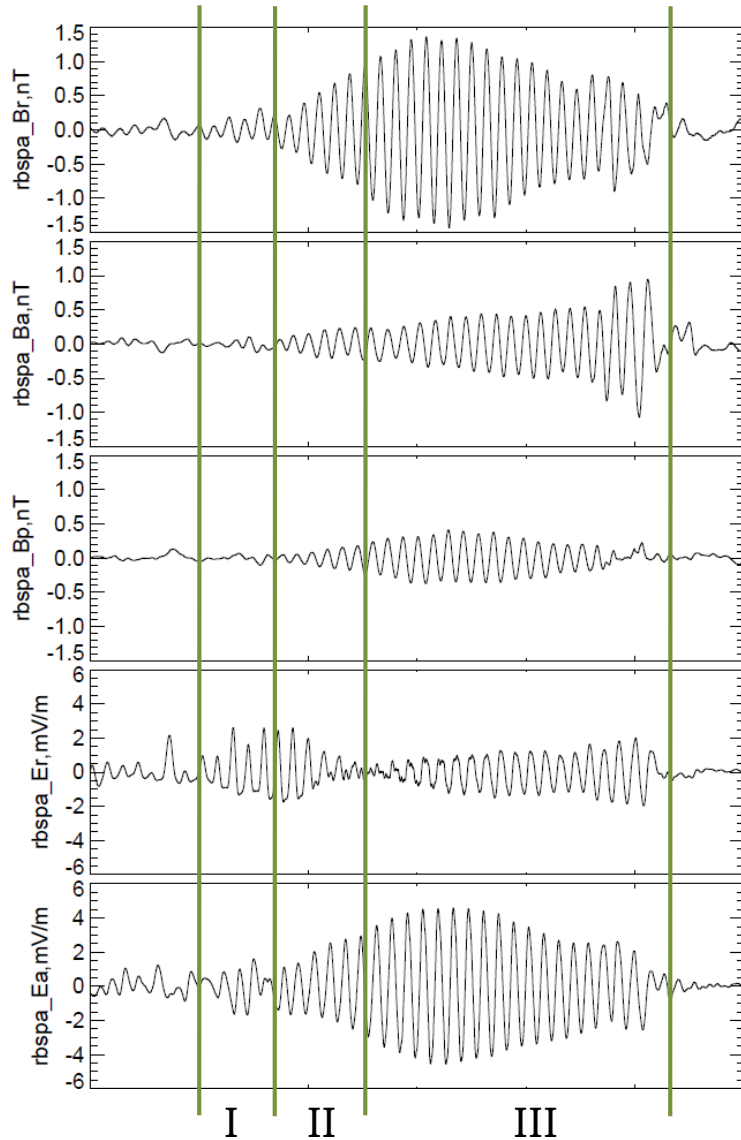


Lei Dai et al. 2013: Poloidal Alfvén wave generated by energetic particles - II

RBSP-A, Oct. 23, 2012, electric field:

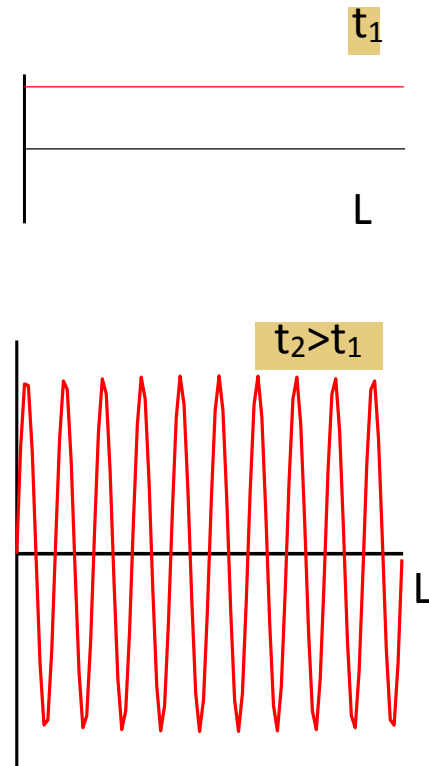


Lei Dai et al. 2013: Poloidal Alfvén wave generated by energetic particles - III



Transformation in space – or in time?

- A newly born (impulse-excited) wave is poloidal, but it evolves into the toroidal due to the phase mixing (because Ω depends on L)
 - Radoski, 1974: box model
 - Leonovich & Mazur, 1998: with field line curvature
 - Klimushkin & Mager, 2004: with field line curvature and instability due to the wave-particle interaction



- However, in this case $\Omega(L)$, but in the Lei Dai's case $\Omega = \text{almost const}(L)$
- Probably, they observed monochromatic wave...
- ... And transformation was in space (due to the curvature), not in time

Related publications

- Leonovich A. S. and Mazur, V. A., A theory of transverse small scale standing Alfvén waves in an axially symmetric magnetosphere, *Planet. Space Sci.*, 41, 697–717, 1993.
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- Leonovich, A. S. and Mazur, V. A., Standing Alfvén waves with $m \gg 1$ in an axisymmetric magnetosphere excited by a non-stationary source, *Ann. Geophysicae*, 16, 914–920, 1998.
- Klimushkin D. Yu., Structure of small-scale standing azimuthal Alfvén waves interacting with high-energy particles in the magnetosphere, *Plasma Phys. Rep.*, V.24, p.956-964, 1998.
- Klimushkin D. Yu., P.N. Mager, K.-H. Glassmeier, Toroidal and poloidal Alfvén waves with arbitrary azimuthal wave numbers in a finite pressure plasma in the Earth's magnetosphere, *Ann.Geophys.*, V. 22, No1, pp. 267-288, 2004.
- D. Yu. Klimushkin, P.N. Mager, The spatio-temporal structure of impulse-generated azimuthally small-scale Alfvén waves interacting with high-energy charged particles in the magnetosphere, *Ann.Geophys.*, V. 22, pp. 1053–1060, 2004.
- Dai L., et al., Excitation of poloidal standing Alfvén waves through drift resonance wave-particle interaction, *Geophys. Res. Lett.*, 40, 4127–4132, doi:10.1002/grl.50800, 2013.
- Cheremnykh O.K., Klimushkin D., Kostarev D.V., On the structure of azimuthally small-scale ULF oscillations of hot space plasma in a curved magnetic field. Modes with continuous spectrum, *Kinematics and physics of celestial bodies*. V.30, P. 209-222, 2014.
- Leonovich A. S., Klimushkin D. Yu., Mager P.N., Experimental evidence for the existence of monochromatic transverse small-scale standing Alfvén waves with spatially dependent polarization, *JGR* (submitted).

Thank you!

